# University of California, Berkeley <br> Department of Mechanical Engineering <br> ME 104, Fall 2013 <br> <br> Midterm Exam 2 Solutions 

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1. (30 points) Consider a particle $X$, of mass $m$, moving along a plane curve $\mathcal{C}$, which is parametrized by its arc length $s$. The unit tangent vector $\mathbf{e}_{t}$ to $\mathcal{C}$ is defined by

$$
\begin{equation*}
\mathbf{e}_{t}=\frac{d \mathbf{r}}{d s} \tag{1}
\end{equation*}
$$

and the principal unit normal vector to $\mathcal{C}$ is

$$
\begin{equation*}
\mathbf{e}_{n}=\rho \frac{d \mathbf{e}_{t}}{d s} \tag{2}
\end{equation*}
$$

where $\rho(>0)$ is the radius of curvature of $\mathcal{C}$.
(a) Show that the velocity and acceleration vectors of $X$ may be expressed as

$$
\begin{equation*}
\mathbf{v}=\dot{s} \mathbf{e}_{t}=v \mathbf{e}_{t}, \quad \mathbf{a}=\dot{v} \mathbf{e}_{t}+\frac{v^{2}}{\rho} \mathbf{e}_{n} \tag{3}
\end{equation*}
$$

(Solution) (10 points)

$$
\begin{gathered}
\mathbf{v}=\dot{\mathbf{r}}=\frac{d \mathbf{r}}{d s} \dot{s}=\dot{s} \mathbf{e}_{t}=v \mathbf{e}_{t} \\
\mathbf{a}=\dot{\mathbf{v}}=\dot{v} \mathbf{e}_{t}+v \dot{\mathbf{e}}_{t}=\dot{v} \mathbf{e}_{t}+v \frac{d \mathbf{e}_{t}}{d s} \dot{s}=\dot{v} \mathbf{e}_{t}+\frac{v^{2}}{\rho} \mathbf{e}_{n}
\end{gathered}
$$

(b) Suppose that a car weighing 3000 lbf is being tested on a circular skidpad of radius 100 ft . If the maximum sustainable constant speed of the car is 37 mph , calculate the acceleration of the car in $g$ 's. Also, calculate the total frictional force that the asphalt surface then exerts on the tires.
(Solution) (10 points)
The car has a constant speed $v=37 \mathrm{mph}$, so

$$
\mathbf{a}=0 \mathbf{e}_{t}+\frac{v^{2}}{\rho} \mathbf{e}_{n}=\frac{v^{2}}{\rho} \mathbf{e}_{n} .
$$

The acceleration of the car is

$$
a_{n}=\frac{v^{2}}{r}=\left(\frac{37 \times 5280}{3600}\right)^{2}\left(\frac{1}{100}\right)=29.4487 \mathrm{ft} / \mathrm{s}^{2}=0.915 \mathrm{~g} .
$$

The total frictional force on the tires is

$$
F=m a_{n}=m \frac{v^{2}}{r}=\left(\frac{3000}{32.2}\right) 29.4487=2740 \mathrm{lbf}
$$

(c) Suppose that a projectile is launched with speed $v_{0}$ at an angle $45^{\circ}$ from the horizontal plane. Neglect air resistance. Draw the free-body diagram of the projectile at the highest point $A$ of its trajectory. Find the velocity vector $\mathbf{v}(A)$ and acceleration vector a $(A)$ of the projectile, and indicate them on your sketch. What are the normal and tangential components of $\mathbf{a}(A)$ ? Express the radius of curvature $\rho(A)$ in terms of $v_{0}$ and $g$.
(Solution) (10 points)
At the highest point $A$, the velocity $\mathbf{v}$ is given by

$$
\mathbf{v}(A)=v_{0 x} \mathbf{i}+0 \mathbf{j}=\frac{v_{0}}{\sqrt{2}} \mathbf{i}
$$

The acceleration is

$$
\mathbf{a}(A)=-g \mathbf{j}=g \mathbf{e}_{n}=\frac{v^{2}}{\rho} \mathbf{e}_{n}
$$

The tangential component of $\mathbf{a}$ is zero. Taking the normal component,

$$
g=\frac{v_{0}^{2}}{2 \rho} \quad \longrightarrow \quad \rho(A)=\frac{v_{0}^{2}}{2 g}
$$


2. (35 points) Recall that the linear impulse of a force $\mathbf{F}$ is defined by

$$
\begin{equation*}
\mathcal{I}=\int_{t_{1}}^{t_{2}} \mathbf{F} d t \tag{4}
\end{equation*}
$$

(a) If $\mathbf{G}$ is the linear momentum of a particle of mass $m$, prove that

$$
\begin{equation*}
\mathcal{I}=\Delta \mathbf{G}=\mathbf{G}_{2}-\mathbf{G}_{1} . \tag{5}
\end{equation*}
$$

(Solution) (5 points)

$$
\mathcal{I}=\int_{t_{1}}^{t_{2}} \mathbf{F} d t=\int_{t_{1}}^{t_{2}} \dot{\mathbf{G}} d t=\int_{t_{1}}^{t_{2}} \frac{d \mathbf{G}}{d t} d t=\int_{\mathbf{G}_{1}}^{\mathbf{G}_{2}} d \mathbf{G}=\mathbf{G}_{2}-\mathbf{G}_{1}=\Delta \mathbf{G}
$$

(b) A simplified model of a ballistic pendulum comprises: (i) a simple pendulum consisting of an inextensible string of length $\ell$ and a bob of mass $M$, which is initially at rest in its stable equilibrium position; and (ii) a bullet of mass $m$ travelling with speed $v$ on a collision course with the bob. Treat the collision as impulsive, and calculate the speed $v^{+}$of the bob and the embedded bullet immediately after impact.
(Solution) (10 points)
Gravitation force is neglected during impact, and the impulsive forces are entirely internal to the system. Thus, the linear momentum is conserved:

$$
\mathbf{G}_{2}=\mathbf{G}_{1} \quad \rightarrow \quad m v \mathbf{j}=(m+M) v^{+} \mathbf{j} \quad \rightarrow \quad v^{+}=\frac{m}{m+M} v
$$

(c) Show that the maximum angle $\theta_{m}$ reached by the pendulum after impact is given by

$$
\begin{equation*}
1-\cos \theta_{m}=\frac{\left(v^{+}\right)^{2}}{2 g \ell} . \tag{6}
\end{equation*}
$$

(Solution) (10 points)
Using conservation of energy between the positions $\theta=\theta_{m}$ and $\theta=0$, we have

$$
\begin{aligned}
T^{+}+V^{+} & =0+V_{m} \\
\frac{1}{2}(M+m)\left(v^{+}\right)^{2}-(M+m) g \ell & =0-(M+m) g \ell \cos \theta_{m}
\end{aligned}
$$

or

$$
\frac{1}{2}(M+m)\left(v^{+}\right)^{2}=(M+m) g \ell\left(1-\cos \theta_{m}\right) .
$$

Rearrange to get

$$
1-\cos \theta_{m}=\frac{\left(v^{+}\right)^{2}}{2 g \ell}
$$

(d) If

$$
\ell=5 \mathrm{~m}, \quad m=0.2 \mathrm{~kg}, \quad M=20 \mathrm{~kg}, \quad \theta_{m}=10^{\circ},
$$

calculate the incoming speed $v$ of the bullet, and express it as a Mach number, using $c=343$ $\mathrm{m} / \mathrm{s}$ for the speed of sound in air at a temperature of $20^{\circ} \mathrm{C}$.
(Solution) (10 points)
From Parts (b) and (c),

$$
\left(\frac{m}{m+M}\right)^{2} v^{2}=2 g \ell\left(1-\cos \theta_{m}\right)
$$

Solve for the speed $v$ and substitute the given values:

$$
\begin{aligned}
v & =\left(\frac{m+M}{m}\right) \sqrt{2 g \ell\left(1-\cos \theta_{m}\right)} \\
& =\left(\frac{0.2+20}{0.2}\right) \sqrt{2 \times 9.81 \times 5 \times\left(1-\cos 10^{\circ}\right)} \\
& =123.3 \mathrm{~m} / \mathrm{s} \quad \longrightarrow \quad M a=\frac{v}{c}=\frac{123.3}{343}=0.36 .
\end{aligned}
$$

3. (35 points) Consider a planet of mass $m$ orbiting the sun (mass $M$ ), regarded as fixed in an inertial frame of reference. Use polar coordinates, and recall that the velocity of the planet is given by

$$
\begin{equation*}
\mathbf{v}=\dot{r} \mathbf{e}_{r}+r \dot{\theta} \mathbf{e}_{\theta} \tag{7}
\end{equation*}
$$

(a) Calculate the angular momentum $\mathbf{H}_{O}$ of the planet about the center of the sun, and deduce that

$$
\begin{equation*}
r^{2} \dot{\theta}=\text { const } .=h . \tag{8}
\end{equation*}
$$

(Solution) (8 points)
The position vector of the planet is $\mathbf{r}=r \mathbf{e}_{r}$. The angular momentum

$$
\begin{aligned}
\mathbf{H}_{O} & =\mathbf{r} \times m \mathbf{v} \\
& =r \mathbf{e}_{r} \times m\left(\dot{r} \mathbf{e}_{r}+r \dot{\theta} \mathbf{e}_{\theta}\right) \\
& =m r^{2} \dot{\theta} \mathbf{k} .
\end{aligned}
$$

The moment $\mathbf{M}_{O}$ about the center of the sun is zero, which means $\dot{\mathbf{H}}_{O}=\mathbf{0}$. Thus, $\mathbf{H}_{O}$ is a constant vector. Then

$$
\frac{\left\|\mathbf{H}_{O}\right\|}{m}=r^{2} \dot{\theta}=\text { const. }=h .
$$

This constant, $h$, is the specific angular momentum of the planet.
(b) Recall that the gravitational potential energy of the planet is $V=-\frac{G M m}{r}$, where $G=6.6734 \times 10^{-11} \mathrm{~m}^{3} / \mathrm{kg} . \mathrm{s}^{2}$, and the mass of the sun is $1.990 \times 10^{30} \mathrm{~kg}$. Deduce that

$$
\begin{equation*}
\frac{1}{2} v^{2}-\frac{G M}{r}=\text { const. }=\frac{E}{m} . \tag{9}
\end{equation*}
$$

(Solution) (7 points)
The Newtonian gravitational force is a conservative force field. Hence, energy is conserved.

$$
T+V=\frac{1}{2} m v^{2}-\frac{G M m}{r}=E=\text { const } .
$$

Divide this equation by the mass $m$ to get (9).
(c) For the planet Mercury, the distances to perihelion and aphelion are

$$
\begin{equation*}
r_{p}=46 \times 10^{6} \mathrm{~km}, \quad r_{a}=69.82 \times 10^{6} \mathrm{~km} \tag{10}
\end{equation*}
$$

If the speed of Mercury at aphelion is $38.86 \times 10^{3} \mathrm{~m} / \mathrm{s}$, calculate its speed at perihelion.
(Solution) (7 points)
The constant specific angular momentum

$$
h=r_{a} v_{a}=r_{p} v_{p} .
$$

Thus,

$$
v_{p}=v_{a} \frac{r_{a}}{r_{p}}=\left(38.86 \times 10^{3} \mathrm{~m} / \mathrm{s}\right) \frac{69.82 \times 10^{6} \mathrm{~km}}{46 \times 10^{6} \mathrm{~km}}=58.98 \times 10^{3} \mathrm{~m} / \mathrm{s}
$$

(d) Using aphelion data, calculate the value of specific orbital energy of Mercury.
(Solution) (7 points)
Directly using Eq. (9), the specific orbital energy

$$
\begin{aligned}
\frac{E}{m} & =\frac{1}{2} v_{a}^{2}-\frac{G M}{r_{a}} \\
& =\frac{1}{2}\left(38.86 \times 10^{3} \mathrm{~m} / \mathrm{s}\right)^{2}-\frac{\left(6.6734 \times 10^{-11} \mathrm{~m}^{3} / \mathrm{kg} \cdot \mathrm{~s}^{2}\right)\left(1.990 \times 10^{30} \mathrm{~kg}\right)}{69.82 \times 10^{9} \mathrm{~m}} \\
& =-1.147 \times 10^{9} \mathrm{~m}^{2} / \mathrm{s}^{2}
\end{aligned}
$$

(e) Where in its orbit does the maximum acceleration of Mercury occur? Calculate its value.
(Solution) (6 points)
Recall the equation of orbital motion $\mathbf{F}=m \mathbf{a}$ :

$$
-\frac{G M m}{r^{2}} \mathbf{e}_{r}=m\left(\ddot{r}-r \dot{\theta^{2}}\right) \mathbf{e}_{r}+m(r \ddot{\theta}+2 \dot{r} \dot{\theta}) \mathbf{e}_{\theta}
$$

The magnitude of the acceleration $a=\|\mathbf{F}\| / m=G M / r^{2}$. It's clear that the maximum acceleration occurs at minimum $r$, which is at perihelion. This maximum value is

$$
\begin{aligned}
a_{\max }=\frac{G M}{r_{p}^{2}} & =\frac{\left(6.6734 \times 10^{-11} \mathrm{~m}^{3} / \mathrm{kg} \cdot \mathrm{~s}^{2}\right)\left(1.990 \times 10^{30} \mathrm{~kg}\right)}{\left(46 \times 10^{9} \mathrm{~m}\right)^{2}} \\
& =6.276 \times 10^{-2} \mathrm{~m} / \mathrm{s}^{2} .
\end{aligned}
$$

Note that

$$
a_{\max } \neq r_{p} \dot{\theta}_{p}^{2}=\frac{v_{p}^{2}}{r_{p}}=7.563 \times 10^{-2} \mathrm{~m} / \mathrm{s}^{2}
$$

because $\ddot{r}_{p}$ is not zero, and $r_{p} \neq \rho=55.4 \times 10^{9} \mathrm{~m}$.

